Using adaptive nested mesh code HydroBox3D for numerical simulation of Type Ia supernovae: merger of carbon-oxygen white dwarf stars, collapse, and non-central explosion

Igor Kulikov, Igor Chernykh, Dmitry Karavaev, Victor Protasov, Alexander Serenko, Vladimir Prigarin, Ivan Ulyanichev ICMMG SB RAS, Novosibirsk

> Alexander Tutukov Institute of Astronomy RAS, Moscow

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The motivation: white dwarf evolution



The motivation: white dwarf evolution (asymmetric explosion)



The mathematical challenges:

- 1. The numerical model construction
- 2. The numerical solver development
- 3. The efficiency parallel implementation

White dwarf expands less than in the Röpke et al. model, so the collision on the far side occurs at higher density and with less geometrical dilution. In the Chicago version, the temperature is sufficient to ignite a detonation that consumes the rest of the star.

Jordan et al. (2008)

The supernova explosion enriches interstellar medium with the elements of life: *O, C, Fe, N, Si, Mg, Ca,...*

The Hydrodynamic Model of White Dwarf

- The Euler hydrodynamics equations
- The gravity
- The stellar equation of state:
 - Ideal gas for low temperature
 - Adiabatic (non)relativistic degenerate gas for high temperature
 - Radiation term
- The carbon burning ¹²C + ¹²C -> ²³Na + p

The numerical methods for hydrodynamics

SPH approach

- Robustness of the algorithm
- Galilean-invariant solution
- Simplicity of implementation
- Flexible geometries of problems
- High accurate gravity solvers

AMR approach

- Approved numerical methods
- No artificial viscosity
- Higher order shock waves
- Resolution of discontinuities
- No suppression of instabilities
- Correct turbulence solution
- Artificial viscosity is parameterized
- Variations of the smoothing length
- The problem of shock wave and discontinuous solutions
- Instabilities suppressed
- The method is not scalable

- The complexity of implementation
- The effects of mesh
- Problem of the minimal mesh
- Not Galilean-invariant solution
- The method is not scalable

Top10 (November 11, 2018) (TFlop/s) (TFlop/s) (kW) System Cores Summit - IBM Power System AC922, IBM POWER9 22C 3.07GHz, NVIDIA 2.397.824 143.500.0 200.794.9 9.783 olta GV100, Dual-rail Mellanox EDR Infiniband , IBM SC/Oak Ridge National Laboratory United St Sierra - IBM Power System S922LC, IBM POWER9 22C 3.1GHz, NVIDIA 1,572,480 94,640.0 125,712.0 7,438 olta GV100, Dual-rail Mes nox EDR Infiniband , IBM / NVIDIA / Mellanox 10E/IN SA/LUNE United States wwwyTaihuLight - Sunway 260C 1.45GHz. 10.649.600 93.014.6 125.435.9 15.371 Sunway , NRCPC brid supercomputers National Supercomputing Center in Wuxi P China sinhe-2A - TH-IVB-FEP Cluster, Int 4,981,760 61,444,5 100,678,7 18,482 2GHz, TH Express-2, Matrix-2000, NUDT National Super Com ter Center in China Daint - Cray XC50 Hz, Aries interconnect , 387,872 21,230.0 27,154.3 2,384 NVIDIA Tesla P10 Cray Inc . Supercom Swiss Nati ung 16C 2.3GHz. Intel Xeon Phi 7250 68C 979.072 20.158.7 41.461.2 7.578 mity - Cray 4GHz interconnect . Cra Inc 54/LANL/SN DOE/ Al Bridging Woud Infrastructure (ABCI) - PRIMERGY CX2570 M4, Xeon 391,680 19,880.0 32,576.6 1,649 Gold 6148 20C 2.4GHz, NVIDIA Tesla V100 SXM2, Infiniband EDR , Fujitsu of Advanced Industrial Science and Technology (AIST) Unstitute SuperMU2-NG - ThinkSystem SD530, Xeon Platinum 8174 24C 3.1GHz, 305.856 19.476.6 26.873.9 Intel Orini-Path , Lenovo iz Rechenzentrum Leib rmany Titan - Cray XK7, Opteron 6274 16C 2.200GHz, Cray Gemini interconnect, 560.640 17.590.0 27.112.5 8.209 NVIDIA K20x, Cray Inc. DOE/SC/Oak Ridge National Laboratory United States Seguoia - BlueGene/Q, Power BQC 16C 1.60 GHz, Custom , IBM 1.572.864 17.173.2 20.132.7 7.890 10 DOE/NNSA/LLNL United States

The original numerical methods



The Parallel Implementation



_mm512_set1_pd
 set value for a vector
_mm512_load_pd
 load a vector from main memory
_mm512_mul_pd
 vector multiply
_mm512_add_pd
 vector summation
_mm512_sub_pd
 vector substitution
_mm512_store_pd
 store a vector to main memory

Main advantages is 302 GFLOPS on Intel Xeon Phi KNL

Main disadvantages – formation of the 8-double elements vector for computing

Pitfalls: associative of cache memory, align of memory, schedule distribution, data dependency

The Sedov explosion



The Evrard collapse





Supernovae explosion

Star formation





Protoplanetary disks and planet formation



CPU



CPU

Many CPU







The Base/Satellite Computing



Shared Memory

The organization of Base Computing





The adjacent for Riemann problem





The mesh reconstruction





The Carbon burning (The Satellite Computing)



The asymmetric explosion of white dwarf (The Base Computing)









Supernovae Ib/Ic/II (ISP RAS Open 2019)



The Base Computing





The Satellite Computing

Credit ESO and Wiki

Conclusion

- A new numerical model of Supernovae la type explosion is created.
- The Base/Satellite computing concept is described.
- The scenarios of a non-central SNIa is modeled.
- The SNIa is non standard problem!!!

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